# Submillimeter Spectrum and Structure of Chloryl Chloride, ClClO<sub>2</sub>

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The submillimeter spectrum of ClClO<sub>2</sub> has been observed between 365 and 398 GHz in a flowing system as a product of the FClO<sub>2</sub> + HCl reaction. Absorptions of three isotopomers,  ${}^{35}Cl^{35}ClO_2$ ,  ${}^{37}Cl^{35}ClO_2$ , and  ${}^{35}Cl^{37}ClO_2$ , have been assigned. Rotational and centrifugal distortion constants have been determined and structural parameters derived. The data are compared with those from a matrix-isolation study, *ab initio* calculations, and related molecules.

## Introduction

Chlorine oxides are among the most important molecules to participate in the catalytic ozone destruction. About a decade ago, a catalytic cycle involving the dimerization of CIO to CIOOCI was proposed as having a key role in the formation of the Antarctic ozone hole. The CIOCIO form of  $Cl_2O_2$  was thought not to contribute to net ozone loss.<sup>1</sup> Initial *ab initio* calculations suggested three  $Cl_2O_2$  isomers, dichlorine peroxide, CIOOCI, chloryl chloride, CICIO<sub>2</sub>, and chlorine chlorite, CIO-CIO, to be thermodynamically stable under stratospheric conditions.<sup>2</sup> More sophisticated calculations indicate that CIOOCI and CICIO<sub>2</sub> differ in their heat of formation by less than the suggested uncertainties of ~8 kJ mol<sup>-1</sup> and are approximately 40 kJ mol<sup>-1</sup> more stable than CIOCIO.<sup>3</sup>

Both ClOOCl and ClClO<sub>2</sub> have been directly observed in the gas phase as relatively long-lived species.<sup>4,5</sup> The molecular structure of the peroxide has been determined from its rotational spectrum,<sup>4</sup> and the harmonic force field has been derived from its IR spectrum and centrifugal distortion constants.<sup>6</sup> The harmonic force field and a preliminary structure of ClClO<sub>2</sub> have been derived from its IR spectra in inert gas matrices<sup>7</sup> and in the gas phase.<sup>5</sup> ClOClO has not been observed directly in the gas phase although evidence has been presented for its formation as a transient product of the ClO self-reaction.<sup>5,8</sup>

More recently ClClO<sub>2</sub> was observed in the ClO self-reaction on ice<sup>9</sup> and in the photochemistry of OClO in and on ice, respectively.<sup>10,11</sup> In a photoionization study on products of the ClO dimerization in the gas phase evidence was found for the formation of another Cl<sub>2</sub>O<sub>2</sub> isomer besides ClOOCl.<sup>12</sup> These findings have raised the question of the importance of the different Cl<sub>2</sub>O<sub>2</sub> isomers in the ClO dimerization and of the role of ClClO<sub>2</sub> in stratospheric chemistry in general.

As part of a program to provide a millimeter and submillimeter spectroscopic data base of potentially important upper atmospheric molecules and to determine molecular properties which may influence chemical behavior, we have studied the rotational spectrum of ClClO<sub>2</sub>. This Letter reports an analysis of the spectra of three isotopomers between 365 and 398 GHz, the spectroscopic constants, and the derived structural parameters. Work is in progress on further measurements of hyperfine patterns to determine the quadrupole coupling constants of both Cl nuclei. Since *ab initio* calculations are often used to



Figure 1. Projection of the ClClO<sub>2</sub> molecule into the *ac*-plane.

determine thermochemical properties of molecules, the structural information reported here will provide an important reference point for such calculations. It is also of interest to compare the structure with that deduced from a matrix IR study since the reliability of the latter is often difficult to assess.

### **Experimental Details**

ClClO<sub>2</sub> was prepared *in situ* by the reaction of FClO<sub>2</sub> with HCl. FClO<sub>2</sub> was prepared from ClF<sub>3</sub> (Spectra Gases) and NaClO<sub>3</sub>. A commercial sample of HCl (Spectra Gases, 99.995%) was used. Conditions for ClClO<sub>2</sub> production were optimized by first observing the UV spectrum of the reaction products in the millimeter sample cell. At ~255 K about 10– 30% of the total pressure was due to ClClO<sub>2</sub> when FClO<sub>2</sub> was mixed with about stoichiometric amounts of HCl in a ~1 m long, 2.5 cm i.d. prereactor at pressures of 10–30 Pa, and the products were expanded through a ~1.5 mm pinhole into the 1 m long, 5 cm i.d. absorption cell. The pumping speed was low to minimize the amount of chemicals used. It was adjusted so that a pressure inside the cell was in the 1 Pa range, suitable for submillimeter spectroscopy.

The spectrometer is essentially identical to that described by Friedl *et al.*<sup>13</sup> The entire region between 378.5 and 395.0 was scanned. Additional measurements were made outside this range in order to extend and confirm the initial assignments of R-branch series having the same  $K_a$ . Lines of unreacted FClO<sub>2</sub> and of OClO, a decomposition product of ClClO<sub>2</sub> and a hydrolysis product of FClO<sub>2</sub>, were eliminated from consideration by comparison with reference spectra and calculated line positions.

## Results

Initial predictions of the rotational spectrum of  $ClClO_2$  were made using the structure from ref 7 together with the quartic distortion constants derived from the harmonic force field. The

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TABLE 1: Spectroscopic Constants<sup>a</sup> (MHz) of Chloryl Chloride, ClClO<sub>2</sub>

parameter	<sup>35</sup> Cl <sup>35</sup> ClO <sub>2</sub>	<sup>37</sup> Cl <sup>35</sup> ClO <sub>2</sub>	<sup>35</sup> Cl <sup>37</sup> ClO <sub>2</sub>	parameter	<sup>35</sup> Cl <sup>35</sup> ClO <sub>2</sub>	<sup>37</sup> Cl <sup>35</sup> ClO <sub>2</sub>	<sup>35</sup> Cl <sup>37</sup> ClO <sub>2</sub>
Α	9450.9218 (98)	9449.8443 (396)	9414.6640 (99)	$\delta_{ m J}  imes 10^3$	0.3842 (4)	0.3537 (13)	0.3754 (6)
В	3588.0617 (62)	3469.5738 (229)	3565.0887 (94)	$\delta_{ m K}  imes 10^3$	6.1348 (43)	5.8911 (130)	6.0322 (58)
С	2787.9468 (79)	2715.9436 (232)	2777.2698 (123)	$\Phi_{ m JK}  imes 10^7$	0.113 (22)	b	b
$\Delta_{ m J}  imes 10^3$	1.7603 (13)	1.6600 (35)	1.7350 (19)	$\Phi_{ m KJ}  imes 10^7$	0.221 (74)	b	b
$\Delta_{ m JK}  imes 10^3$	6.5029 (38)	6.2313 (125)	6.4457 (46)	$\Phi_{ m K}  imes 10^7$	-0.283 (63)	b	b
$\Delta_{\rm K}  imes 10^3$	-1.2229 (147)	-0.8188 (463)	-1.2450 (147)	$\phi_{ m K}  imes 10^7$	1.837 (223)	b	b

<sup>*a*</sup> Numbers in parentheses are 1 standard deviation in units of the least significant figures. Watson's A reduction in the representation I<sup>r</sup> was used. <sup>*b*</sup> Common constants for all three isotopomers.

TABLE 2: Structural Parameters (pm, deg) of ClClO<sub>2</sub> and Related Molecules

	ClClO <sub>2</sub> , exptl		ClClO <sub>2</sub> , ab initio				FClO <sub>2</sub> <sup>e</sup>	OClO <sup>f</sup>
	gas phase <sup>a</sup>	matrix <sup>b</sup>	MP2/6-31G*c	$MP2/TZ2P^d$	MP2/TZ2Pf <sup>d</sup>	CCSD(T)/TZ2Pd	exptl, gas phase	
$r(\text{XCl})$ $r(\text{ClO})$ $\angle(\text{XClO})$ $\angle(\text{OClO})$	219.78 (23) 143.60 (12) 104.06 (14) 115.05 (15)	222. (6) 144.0 (5) 103.5 (10) 116.0 (5)	245.9 146.9 104.2 116.7	233.4 145.1 103.9 115.6	226.9 144.2 103.8 115.0	227.9 146.8 104.1 114.9	169.43 142.03 101.81 115.25	146.98 117.40

<sup>*a*</sup> r<sub>o</sub> structure, this work. <sup>*b*</sup> Reference 7. <sup>*c*</sup> Reference 2. <sup>*d*</sup> Reference 3. <sup>*e*</sup> r<sub>0</sub> structure, ref 14. <sup>*f*</sup> r<sub>e</sub> structure, ref 15.

pyramidal ClClO<sub>2</sub> is an asymmetric prolate top,  $\kappa = -0.7598$  for <sup>35</sup>Cl<sup>35</sup>ClO<sub>2</sub>. It has  $C_s$  symmetry with dipole components along the *a*- and *c*-axes, as shown in Figure 1. In the region searched, the strongest features in the spectrum are expected to be high  $K_a$  *c*-type R-branches. Because  $A - (B + C)/2 \approx B + C$ , each R-branch has its origin in the vicinity of the third member of the branch with one lower  $K_a$ . At high  $K_a$  this leads to some clustering of the stronger transitions and facilitates the identification of different series once one has been identified.

Initially a series of strong, partially resolved doublets with spacings of approximately the estimated value of B + C was identified and eventually extended to six consecutive members. Comparison with predicted spectra and some trial fits quickly led to the correct assignment of the J'' = 24-29 members of the  $K_a'' = 16$  rR branch and the identification of further transitions. Because of higher quantum numbers, large uncertainties, and a possibly small value for  $\mu_a$ , the *a*-dipole transitions have not yet been unambiguously assigned. Most of the observed lines showed small effects of quadrupole coupling due to the two Cl nuclei, generally in the form of doublets. The centers of the patterns were used in fitting the spectra. The quadrupole analysis will appear in a subsequent paper. Rotational, quartic, and some sextic centrifugal distortion constants have been determined. Because not all of the sextic distortion constants were determinable by the present data, the final values of the molecular constants may lie outside the quoted uncertainties. For the rotational and quartic distortion constants as well as for the derived structure these effects are expected to be small.

A preliminary structure was derived from the rotational constants of  ${}^{35}Cl^{35}ClO_2$  and an estimate of r(ClO) in order to predict rotational constants of  ${}^{37}Cl^{35}ClO_2$  and  ${}^{35}Cl^{37}ClO_2$ . Quartic distortion constants were estimated from the harmonic force field.<sup>7</sup> Sextic constants were taken from the main isotopomer. Some lines were found for both isotopomers within a few megahertz of the predictions enabeling further transitions to be assigned. Although the number of observed transitions was smaller for the less abundant isotopomers than for  ${}^{35}Cl^{35}ClO_2$ , rotational and quartic distortion constants could be determined precisely.

The spectroscopic constants from a combined fit of all three isotopomers are given in Table 1. The principal planar moments were used to derive a ground state effective,  $r_0$ , structure which is given in Table 2. The uncertainties quoted are those derived from the standard deviation of the fit. Because of vibrational effects inherent in  $r_0$  structure determinations, parameters determined from a different set of isotopomers may lie outside the quoted uncertainties. Matrix and *ab initio* data as well as those for the related  $FCIO_2$  and OCIO molecules are also included.

### Discussion

The structural parameters obtained in the matrix study agree well with the present, more precise gas phase values (Table 2). At the MP2/6-31G\* level the ClCl distance is poorly predicted, and the remaining parameters are in reasonable agreement with the experiment. MP2 calculations employing much larger basis sets yielded more reasonable values for r(ClCl), with the other values also in better agreement. The highest level coupled cluster structure calculation published obtained a better ClCl distance than the MP2 calculation with the same basis set with the remaining constants of similar quality.

It should be mentioned that the MP2/6-31G\* method is frequently used to estimate thermodynamic properties. It predicted ClClO<sub>2</sub> to be higher in energy than ClOOCl by about 100 kJ mol<sup>-1,2</sup> whereas calculations with larger basis sets predict the difference to be *ca*.  $4 \pm 8$  kJ mol<sup>-1,3</sup>

To our knowledge, ClClO<sub>2</sub> constitutes the first determination of a bonding ClCl distance with one of the Cl atoms bonded to more than one atom. For both ClClO<sub>2</sub> and FClO<sub>2</sub> the XCl bond is longer than that in the respective XCl molecule. However, for FClO<sub>2</sub> the difference is about 7 pm,<sup>14,16</sup> while for ClClO<sub>2</sub> it is ~20 pm.<sup>17</sup> The larger XClO bond angle in ClClO<sub>2</sub> compared to FClO<sub>2</sub> may be due to the steric effect of the larger Cl atom.

As has been pointed out earlier,<sup>7</sup> the substitution of an electronegative atom at the Cl atom of OCIO lowers the electron density in antibonding orbitals, thus shortening the ClO bond length. Since F is more electronegative than Cl, its effect is larger. The substitution also reduces the OCIO angle, but the electronegativity seems to have little effect on the magnitude of reduction, since the OCIO angles are the same in ClClO<sub>2</sub> and FClO<sub>2</sub> within experimental uncertainties.

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## **References and Notes**

(1) Molina, L. T.; Molina, M. J. J. Phys. Chem. 1987, 91, 433.

(2) McGrath, M. P.; Clemitshaw, K. C.; Rowland, F. S.; Hehre, W. J. J. Phys. Chem. **1990**, *94*, 6126.

(3) Lee, T. J.; McMichael Rohlfing, C.; Rice, J. E. J. Chem. Phys. 1992, 97, 6593.

(4) Birk, M.; Friedl, R. R.; Cohen, E. A.; Pickett, H. M.; Sander, S. P. J. Chem. Phys. **1989**, *91*, 6588.

(5) Müller, H. S. P.; Willner, H. Ber. Bunsen-Ges. Phys. Chem. 1992, 96, 427.

(6) Jacobs, J.; Kronberg, M.; Müller, H. S. P.; Willner, H. J. Am. Chem. Soc. 1994, 116, 1106.

(7) Müller, H. S. P.; Willner, H. Inorg. Chem. 1992, 31, 2527.

(8) Andrews, L.; Raymond, J. J. Chem. Phys. 1971, 55, 3087.

(9) Kirchner, U.; Liesner, M.; Benter, T.; Schindler, R. N. Tropospheric Oxidation Mechanisms. *Proceedings of the EC/EUROTRAC/GDCh Meeting*, Sept 20–22, 1994, Leipzig, Germany; Becker, K. H., Ed.; Air Pollution Research Report 54.

(10) Pursell, C. J.; Conyers, J.; Denison, C. J. Phys. Chem. 1996, 100, 15450.

(11) Graham, J. D.; Roberts, J. T.; Anderson, L. D.; Grassian, V. H. J. Phys. Chem. **1996**, 100, 19551.

(12) Schwell, M.; Jochims, H.-W.; Wassermann, B.; Rockland, U.; Flesch, R.; Rühl, E. J. Phys. Chem. **1996**, 100, 10070.

(13) Friedl, R. R.; Birk, M.; Oh, J. J.; Cohen, E. A. J. Mol. Spectrosc. **1995**, *170*, 383.

(14) Robiette, A. G.; Parent, C. R.; Gerry, M. C. L. J. Mol. Spectrosc. 1981, 86, 455.

(15) Miyazaki, K.; Tanoura, M.; Tanaka, K.; Tanaka, T. J. Mol. Spectrosc. 1986, 116, 435.

 (16) Willis, R. E., Jr.; Clark, W. W., III. J. Chem. Phys. 1980, 72, 4946.
 (17) Huber, K. P.; Herzberg, G. Molecular Spectra and Molecular Structure; Vol IV: Constants of Diatomic Molecules; Van Nostrand: New York, 1979.